

Sergei Kuksin

# On the Zakharov-L'vov stochastic model for wave turbulence

based on a joint work with Andrey Dymov (Moscow)

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## §1. The setting

Consider the modified NLS equation

$$\frac{\partial}{\partial t}u + i\Delta u - i\nu (|u|^2 - \|u\|^2)u = 0,$$

$$\Delta = (2\pi)^{-2} \sum_{j=1}^d (\partial^2 / \partial x_j^2), \quad x \in \mathbb{T}_L^d = \mathbb{R}^d / (L\mathbb{Z}^d),$$

where  $d \geq 2$ ,  $L \geq 1$  and  $\nu \in (0, 1]$ . This is a hamiltonian PDE, obtained by modifying the standard NLS equation by another hamiltonian equation  $\frac{\partial}{\partial t}u = -i\nu \|u\|^2 u$ , whose flow commutes with that of NLS. This is a rather innocent modification.

Denote by  $H$  the space  $L_2(\mathbb{T}_L^d; \mathbb{C})$ , given the normalised  $L_2$ -norm

$$\|u\|^2 = L^{-d} \Re \int |u|^2 dx; \quad \text{so } \|\mathbf{1}\| = 1.$$

We write solutions  $u$  as  $u(t, x)$  or as  $u(t) \in H$ . Pass to the slow time  $\tau = \nu t$ :

$$\dot{u} + i\nu^{-1} \Delta u - i(|u|^2 - \|u\|^2)u = 0, \quad \dot{u} = (\partial/\partial\tau)u(\tau, x), \quad x \in \mathbb{T}_L^d.$$

From now on I will use the time  $\tau$ .

The objective is to study solutions when  $\nu \rightarrow 0$  and  $L \rightarrow \infty$ .

We write the Fourier series for  $u(x)$  as

$$u(x) = L^{-d/2} \sum_{s \in \mathbb{Z}_L^d} v_s e^{2\pi i s \cdot x}, \quad \mathbb{Z}_L^d = L^{-1} \mathbb{Z}^d,$$

where  $v_s = L^{-d/2} \int_{\mathbb{T}_L^d} u(x) e^{-2\pi i s \cdot x} dx$ .

When studying the equation, people talk about “pumping the energy to low modes and dissipating it in high modes”. To make this rigorous, Zakharov-L’vov in 1975 suggested to consider the NLS equation, dumped by a (hyper)viscosity and driven by a random force:

$$(1) \quad \begin{aligned} \dot{u} + i\nu^{-1} \Delta u - i\rho (|u|^2 - \|u\|^2)u &= -(-\Delta + 1)^{r_*} u + \dot{\eta}^\omega(\tau, x), \\ \eta^\omega(\tau, x) &= L^{-d/2} \sum_s b_s \beta_s^\omega(\tau) e^{2\pi i s \cdot x}. \end{aligned}$$

Here  $r_* > 0$ ,  $\rho \geq 1$  is an additional constant, needed later,  $\{\beta_s(\tau), s \in \mathbb{Z}_L^d\}$  are standard independent complex Wiener processes, the constants  $b_s > 0$  are defined for all  $s \in \mathbb{R}^d$  and fast decay when  $|s| \rightarrow \infty$ .

Denoting  $B = L^{-d} \sum_s b_s^2$  we obtain the balance of energy for solutions of (1):

$$\mathbb{E}\|u(\tau)\|^2 + 2\mathbb{E} \int_0^\tau \|(-\Delta + 1)^{r_*} u(s)\|^2 ds = \mathbb{E}\|u(0)\|^2 + 2B\tau.$$

So the quantity  $\mathbb{E}\|u(\tau)\|^2$  – the averaged “energy per volume” of a solution  $u$  – is order one, uniformly in  $L$ , how this should be.

Passing to the Fourier presentation, we write eq. (1) as

$$\dot{v}_s - i\nu^{-1}|s|^2 v_s + \gamma_s v_s = i\rho L^{-d} \sum_{1,2} \delta'_{3s}{}^{12} v_1 v_2 \bar{v}_3 + b_s \dot{\beta}_s, \quad s \in \mathbb{Z}_L^d,$$

where  $\gamma_s = (1 + |s|^2)^{r_*}$  and

$$\delta'_{3s}{}^{12} = \begin{cases} 1, & \text{if } s_1 + s_2 = s_3 + s \text{ and } \{s_1, s_2\} \neq \{s_3, s\}, \\ 0, & \text{otherwise.} \end{cases}$$

Using interaction representation  $v_s = \exp(i\nu^{-1}\tau|s|^2) a_s$  we write equations for  $v_s$  as

$$\begin{aligned} \dot{a}_s + \gamma_s a_s &= i\rho \mathcal{Y}_s(a; \nu^{-1}\tau) + b_s \dot{\beta}_s, \quad s \in \mathbb{Z}_L^d, \\ (2) \quad \mathcal{Y}_s(a; t) &= L^{-d} \sum_{1,2} \delta'_{3s}{}^{12} a_1 a_2 \bar{a}_3 e^{it\omega_{3s}^{12}}, \\ \omega_{3s}^{12} &= |s_1|^2 + |s_2|^2 - |s_3|^2 - |s|^2 = -2(s_1 - s) \cdot (s_2 - s). \end{aligned}$$

The *energy spectrum* of a solution  $u(\tau)$  is the function

$$\mathbb{Z}_L^d \ni s \mapsto n_s(\tau) = n_s^{L,\nu}(\tau) = \mathbb{E}|v_s(\tau)|^2 = \mathbb{E}|a_s(\tau)|^2.$$

Traditionally the function  $n_s$  is in the center of attention. We wish to study the solutions of (1) and their energy spectra  $n_s$  when

$$\nu \rightarrow 0, \quad L \rightarrow \infty.$$

Exact meaning of this assumption is not clear. Below we specify it as follows:

$$\begin{aligned} \nu \rightarrow 0 \text{ and } L \geq \nu^{-2-\epsilon} \text{ for some } \epsilon > 0, \\ \text{or first } L \rightarrow \infty \text{ and next } \nu \rightarrow 0. \end{aligned}$$

## §2. Solutions as formal series in $\rho$ .

Consider the equations with the initial condition

$$u(-T) = 0, \quad 0 < T \leq +\infty,$$

and write the solution  $a_s$  as formal series in  $\rho$ :

$$a_s = a_s^{(0)} + \rho a_s^{(1)} + \dots$$

Substituting this decomposition in the  $a$ -equation (2), we see that

$$\dot{a}_s^{(0)}(\tau) + \gamma_s a_s^{(0)}(\tau) = b_s \beta_s(\tau), \quad s \in \mathbb{Z}_L^d.$$

So the processes  $a_s^{(0)}$  are independent Ornstein–Uhlenbeck processes:

$$a_s^{(0)}(\tau) = b_s \int_{-T}^{\tau} e^{-\gamma_s(\tau-l)} d\beta_s(l),$$

while  $a^{(1)}$  satisfies

$$\dot{a}_s^{(1)}(\tau) + \gamma_s a_s^{(1)}(\tau) = i\mathcal{Y}_s(a^{(0)}(\tau); \nu^{-1}\tau), \quad \tau > -T,$$

so

$$a_s^{(1)}(\tau) = i \int_{-T}^{\tau} e^{-\gamma_s(\tau-l)} \mathcal{Y}_s(a^0(l); \nu^{-1}l) dl.$$

That is,  $a_s^{(1)}(\tau)$  is a Wiener chaos of third order. Similar, for  $n \geq 2$ ,

$$a_s^{(n)}(\tau) = i \int_{-T}^{\tau} \sum_{n_1+n_2+n_3=n-1} e^{-\gamma_s(\tau-l)} \mathcal{Y}_s(a_1^{(n_1)}(l), a_2^{(n_2)}(l), a_3^{(n_3)}(l); \nu^{-1}l) dl,$$

is a Wiener chaos of order  $2n + 1$ .

**QUASISOLUTIONS.** The NLS equation is a model which is used to describe various small-amplitude nonlinear processes, neglecting the terms, cubic in the amplitude. So what has real physical meaning rather is not itself a solution  $a_s(\tau)$  of the  $a$ -equation (2), but its quadratic in  $\rho$  part. In the notation above this is :

$$A_s(\tau) = a_s^0(\tau) + \rho a_s^1(\tau) + \rho^2 a_s^2(\tau).$$

We call the the process  $A = \{A_s(\tau)\}$  the QUASISOLUTION.

Consider the energy spectrum of  $A$ ,

$$N_s(\tau) = \mathbb{E}|A_s(\tau)|^2.$$

**QUESTION:** How  $N_s$  behaves when  $\nu \rightarrow 0$  and  $L \rightarrow \infty$ ,  $L \gg \nu^{-2}$  ?

Let us write  $N_s$  as series in  $\rho$ :

$$N_s(\tau) = n_s^0(\tau) + \rho n_s^1(\tau) + \rho^2 n_s^2(\tau) + \rho^3 n_s^3(\tau) + \rho^4 n_s^4(\tau).$$

Here  $n_s^0 = \mathbb{E}|a_s^0|^2$  is a quantity of order 1,

$$n_s^1 = 2\Re\mathbb{E}a_s^0\bar{a}_s^1 = 0, \quad n_s^2 = \mathbb{E}|a_s^1|^2 + 2\Re\mathbb{E}a_s^0\bar{a}_s^2, \text{ etc.}$$

**CALCULATION:** if  $\nu \ll 1$  and  $L \gg \nu^{-2}$ , then

$$n_s^2 \sim \nu, \quad n_s^3, n_s^4 \lesssim \nu^2.$$

So the right scaling for  $\rho$  is  $\rho \sim \nu^{-1/2}$ . Accordingly let us take  $\rho$  in the form

$$\rho = \sqrt{\varepsilon} \nu^{-1/2}, \quad \varepsilon \in (0, 1].$$

### §3. Wave kinetic equation

For a real function  $s \mapsto x_s$  on  $\mathbb{R}^d$  let us consider the Cubic Wave Kinetic Integral

$$K_s(x.) = 2\pi\gamma_s \int_{\Sigma_s} \frac{ds_1 ds_2 |_{\Sigma_s} x_1 x_2 x_3 x_s}{\sqrt{|s_1 - s|^2 + |s_2 - s|^2}} \left( \frac{1}{\gamma_s x_s} + \frac{1}{\gamma_3 x_3} - \frac{1}{\gamma_1 x_1} - \frac{1}{\gamma_2 x_2} \right).$$

Here  $x_j = x_{s_j}$ ,  $j = 1, 2, 3$ , we substitute  $s_3 = s_1 + s_2 - s$ ,

$$\Sigma_s = \{(s_1, s_2) : (s_1 - s) \cdot (s_2 - s) = 0\},$$

and  $ds_1 ds_2 |_{\Sigma_s}$  is the microcanonical measure on  $\Sigma$  (the volume in  $\mathbb{R}^{2d}$ , restricted to  $\Sigma$ ).

**FACT:** the Wave Kinetic operator  $x_s \rightarrow K_s(x.)$  is well defined and “good”.

Consider the Wave Kinetic Equation:

$$(WKE) \quad \dot{m}_s(\tau) = -2\gamma_s m_s(\tau) + 2b_s^2 + \varepsilon K_s(m.(\tau)), \quad s \in \mathbb{R}^d.$$

For small  $\varepsilon$  this is a good equation. It has a unique solution, equal 0 at  $-T$ . Let us denote it  $\{n_s^*(\tau)\}$ .

**Theorem.** Let  $\rho = \sqrt{\varepsilon} \nu^{-1/2}$ , where  $\varepsilon$  is a small constant. Then the energy spectrum  $N_s(\tau)$  is close to the solution  $n_s^*(\tau)$  of (WKE):

$$\|n_s^*(\tau) - N_s(\tau)\| \leq C\varepsilon^2 \quad \forall \tau \geq -T.$$

The solution  $n_s^*(\tau)$  can be written as

$$n_s^*(\tau) = n_s^{*0}(\tau) + \varepsilon n_s^{*1}(\tau) + O(\varepsilon^2),$$

where  $n_s^{*0}(\tau) \sim 1$  solves the linear equation  $(WKE)_{\varepsilon=0}$ , and  $\varepsilon n_s^{*1}(\tau)$  is the nonlinear part of the solution.

Note that  $C\varepsilon^2 \ll |\varepsilon n_s^{*1}(\tau)|$  for small  $\varepsilon$ .

**Remark.** If  $\rho = \sqrt{\varepsilon} \nu^{-1/2}$ , then in the original fast time  $t$  the equation reads:

$$\frac{\partial}{\partial t} u + i\Delta u - i\sqrt{\nu} \sqrt{\varepsilon} (|u|^2 - \|u\|^2) u = -\nu(-\Delta + 1)^{r_*} u + \sqrt{\nu} \dot{\eta}^\omega(\tau, x), \quad \|u(t)\| \sim 1.$$

That is,

1) the time-scale which we use to pass to the kinetic limit is  $\tau = \nu t$ , so the time needed to arrive at the limiting kinetic regime is  $t \sim \nu^{-1}$ ;

2) the coefficient in front of the nonlinearity is

$$\nu^{1/2}.$$

## §4. Higher order in $\rho$ decompositions

Write the solution in  $a$ -presentation as formal series in  $\rho$ :

$$a = a^{(0)} + \rho a^{(1)} + \dots,$$

and accordingly write its energy spectrum as

$$(3) \quad n_s(\tau) = n_s^0(\tau) + \rho n_s^1(\tau) + \rho^2 n_s^2(\tau) + \dots$$

Since

$$n_s^2 \sim \nu, \quad n_s^3, n_s^4 \lesssim \nu^2,$$

it is natural to assume that

$$n_s^k \lesssim \nu^{k/2} \quad \text{for all } \nu \text{ and all } k.$$

If so, then scaling as before  $\rho = \sqrt{\varepsilon} \nu^{-1/2}$ , we would make (3) a nice asymptotical series in  $\varepsilon$ .

But this is WRONG:

**Theorem.** 1) For each  $k$  we have

$$n_s^k \leq C_s^\#(k) \max(\nu^{\lceil k/2 \rceil}, \nu^d),$$

where  $\lceil k/2 \rceil$  – the smallest integer which is  $\geq k/2$ .

2) Moreover, if  $k > 2d$ , then the sum of the integrals which makes the term  $n_s^k$  contains integrals of order  $\sim \nu^d \gg \nu^{\lceil k/2 \rceil}$ .

The integrals of order  $\nu^d$  do not cancel each other. So for big  $k$

$$n_s^k \sim \nu^d, \quad \text{NOT} \quad n_s^k \sim \nu^{\lceil k/2 \rceil}.$$

Then the series

$$n_s(\tau) = n_s^0(\tau) + \rho n_s^1(\tau) + \rho^2 n_s^2(\tau) + \dots$$

with the right scaling  $\rho = \sqrt{\varepsilon} \nu^{-1/2}$  IS NOT an asymptotical series since

$$\rho^k n_s^k(\tau) > \varepsilon^{k/2} \nu^{d-k/2},$$

which is very big for  $k > 2d$  and  $\nu \ll 1$ .